

Study of tropospheric scintillation effects in Ku-band frequency for satellite communication system

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ABSTRACT

Scintillation is a rapid fluctuation of electromagnetic waves in terms of phase and amplitude due to a small-scale inconsistency in the transmission path (or paths) with time. Scintillation exists continuously throughout a day whether during raining or clear sky conditions. The raw signal data need to exclude other propagations factors that include signal fluctuations to further understand the scintillation studies. This paper presents the analysis of tropospheric scintillation data from January 2016 till December 2016 at Ku-band frequency of 12.202 GHz beacon signal. The experimental data from MEASAT 3B were collected and analyzed to see the effect of tropospheric scintillation. The elevation angle of the dish antenna is 77.45°. The highlighted objectives are to analyze the scintillation data at Ku-band, and to compare and validate the results with other scintillation models. The result shows that the stipulated scintillation analysis has higher amplitude, which is 0.73 dB compared to other scintillation analysis which has lower scintillation amplitude: 0.45 dB (Karasawa), 0.42 dB (ITU-R), 0.4 dB (Nadirah & Rafiqul), 0.42 dB (Van De Kamp), and 0.11 dB (Anthony & Mandeep).

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1. INTRODUCTION

Satellite is the main element in most of telecommunication system around the globe. Data information are exchanged and received in various forms by earth stations that are transmitted by a microwave repeater station or also known as communications satellite [1]. The significance of communication satellite is that it is basically utilized for the business and government services purposes such as data transfer services, voice services and etc. The communication satellite is also used for obtaining the corporate assets, sending emails, browse internet and it can likewise be utilized to communicate through a high quality voice compression without taking into account the location around the world [2].

There are several propagation effects which are caused by the tropospheric layer: attenuation due to rain, depolarization, gas absorption and scintillation caused by the atmospheric turbulence [3]. Although in the communication systems, rain has the main impact, scintillation now has become crucial for low-margin systems at high frequencies and low elevation angles [3]. Improperly titled paper may never reach the audience for which it was intended, so be specific. The tropospheric scintillation effects are important to be observed especially in broadcasting and telecommunication services which use the Ku-band frequency in the satellite communication links [4].

Scintillation is one of the events which lead to signal degradation through the earth's atmosphere. There are two types of scintillations: ionospheric scintillation and tropospheric scintillation [2]. This paper

emphasizes on tropospheric scintillation because the frequency focused here is at Ku-band in tropical climate. Several scintillation models have been developed and studied by other researchers in the past, however most of the models can only be used in the four seasons (winter, spring, summer and autumn) countries and cannot be applied in the tropical countries such as Malaysia, Singapore, Thailand, Brazil, and etc. Furthermore, previous scintillation studies' data were collected from non-tropical countries such as US, Finland, Germany, United Kingdom and many more. Hence, the scintillation models cannot be used in tropical countries due to the different pattern of climate compared to the four-season countries. The tropical countries consist of high humidity, uniform temperature and heavy rainfall. There are only a few research activities conducted on scintillation in tropical countries and data collected in Malaysia does not fit with any existing models [5-13]. The scintillation models must be evaluated based on the tropical countries data collection and measurement. The scintillation data for tropical climate countries are very limited typically for high elevation angle. Therefore, it is very important to study the scintillation effects based on the data measured in the tropical countries.

2. LITERATURE REVIEW

Scintillation can be found in ionosphere and troposphere layers. In the ionosphere layer, the irregularities of electron that affect frequencies up to 6 GHz are called ionospheric scintillation. Whereas, the scintillations which affect the troposphere layer for the frequencies band above 3 GHz due to the refractive index irregularities is called tropospheric scintillation. The research aim is to study the tropospheric scintillation which affects the Ku-band. This type of scintillation is also caused by temperature inversion layers and high humidity gradients [4]. With the variety of local climate, the impact of tropospheric is seasonally dependent. The structure of the refractive index in the troposphere can be regarded horizontally layered. Furthermore, with the change in altitude, the variations seem to be appeared as thin layers [4]. Moreover, scintillation conditions affect more significantly to slant path at low elevation angle which is highly slanted to the layer formation. The troposphere common properties of the refractive index are widely known. The index of refraction or the atmospheric radio refractive index, n , is a function of pressure, water vapour content, and temperature [4]. The features of the refractive index are normally described in terms of N units, for n is very close to 1. It is defined as

$$N = (n - 1) \times 10^6 = \frac{77.6}{T} \left(p + 4810 \frac{e}{T} \right) \quad (1)$$

where the atmospheric pressure in millibars (mb) is expressed as p ; Temperature, in degrees K is T ; and e is the water vapour pressure in mb. In (2), the first expression is known as the "dry term"

$$N_{dry} = p \frac{77.6}{T} \quad (2)$$

while the "wet term" is the second expression

$$N_{wet} = 3.732 \times 10^5 \frac{e}{T^2} \quad (3)$$

For frequencies up to 100 GHz the accuracy of Equations (2) to (4) are within 0.5% [4]. The refractivity of the long-term dependence can be represented in an exponential form,

$$N = 315 e^{\frac{-h}{7.36}} \quad (4)$$

where the altitude, in km is h . This estimation is applicable for altitude up to about 15 km [14].

3. RESEARCH METHODOLOGY

Ku-Band signal was transmitted at 12.202 GHz frequency from the MEASAT 3 B at longitude (91.5 East) with vertical polarization that provides more than 100 channels which covers Southeast Asia including Malaysia and Indonesia as well as Australia. Table 1 shows the satellite specifications for ASTRO Broadcast Network System. The satellite raw data consists of 12 months period from January 2016 till December 2016 in order to study the scintillation effect. The raw beacon data consist of scintillation, rain and other propagation data. Therefore, in order to separate the scintillation data from other data, a technique is applied by using filtering in MATLAB. Figure 1 illustrates the steps to acquire the scintillation amplitude. Cumulative Distribution Function (CDF) of the scintillation data is obtained by the division of

the scintillation amplitude into two parts: scintillation fades which is the negative signal amplitude and enhancements which is the positive signal amplitude. Next, the CDF of scintillation fades are calculated monthly and annually. Figure 2 illustrates on how to acquire the monthly and annual cumulative distribution function of scintillation fades.

Table 1. Satellite specifications

Satellite	MEASAT 3B
Site	ASTRO Broadcast Network System, Bukit Jalil, Kuala Lumpur, Malaysia
Ground station longitude	101.70
Ground station latitude	3.15
Satellite position	91.517° E
Elevation angle	77.45°
Azimuth	253.02°
Polarization	Vertical
Frequency	12.202 GHz
Antenna diameter	13 m

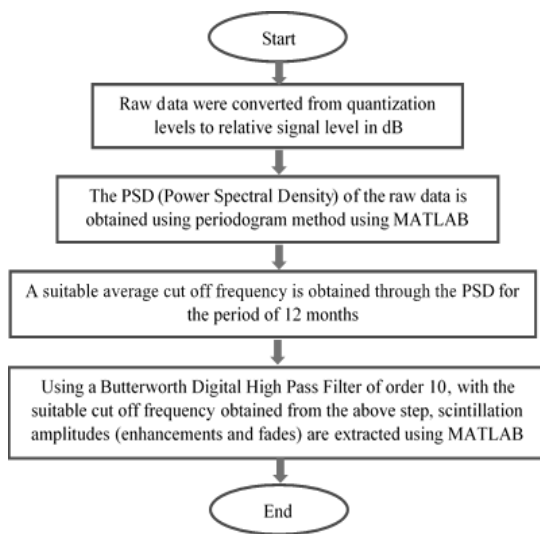


Figure 1. Flowchart of scintillation data extraction from raw data

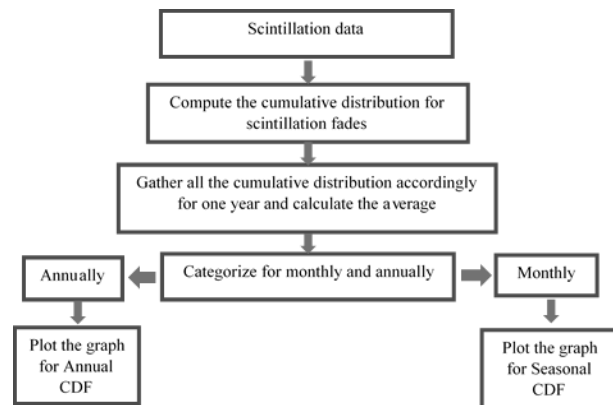


Figure 2. Flowchart to analyze monthly and annual CDF

4. RESULTS AND ANALYSIS

Tropospheric scintillation leads to random scintillation fades and scintillation enhancement of the received signal at a lower elevation angles [14]. Majority of the tropospheric scintillation studies that have been focused on low elevation angles and high frequencies due to the fact that scintillation amplitude is highly observed in this case. However, for the case of high elevation angles and high frequencies only a few studies have been done. This research was done on the basis of high elevation angle and high frequency and is similar to the research in [5-13]. The differences between those researches are the location, elevation angle, frequency, the satellite used and the data sampling as tabulated in Table 2. From Table 2, it can be deduced that this paper, has the same elevation angle as in IIUM but is higher than the one from USM. The humidity and the temperature for each research are different due to the distinct locations. Therefore, due to the difference in humidity and temperature, the scintillation effects will vary for each research [15-16].

Table 2. Research comparison in Malaysia

Research	Location	Elevation Angle (°)	Type of Antenna / Diameter (m)	Frequency (GHz)	Data Sampling (s)	Satellite
USM	Penang, Malaysia	40.1	Cassegrain Antenna / 2.4	12.255	1.0	Superbird C
IIUM	Gombak, Kuala Lumpur, Malaysia	77.5	Dish Antenna / 2.4	10.982	0.1	MEASAT-3
ASTRO	Bukit Jalil, Kuala Lumpur, Malaysia	77.5	Dish Antenna / 13	12.2015	0.05	MEASAT-3B

4.1. Extraction of scintillation data

In order to eliminate other propagation factors that lead to data fluctuations, the raw measured data need to be studied and processed [17]. This is due to the fact that other propagation factors introduce fluctuations at a lower rate than those caused by scintillation [4]. Therefore, the raw data obtained from ASTRO need to go through a high pass filter with an appropriate cut off frequency to observe the scintillation signal level. The data obtained were already divided into two: rain and clear sky. For the purpose of this research, clear sky data are used to discard any data that contribute to rain attenuation. The raw data obtained were recorded in terms of attenuation in dB; hence, the data need to be subtracted by 75 dB to obtain the original received signal level. Figure 3 illustrates the received signal level of measured raw data versus number of data for 22 February 2016. The data are usually represented in terms of percentage of time [16] but the number of samples and sampling time vary for each day for the raw data received from ASTRO thus, the data are represented in terms of number of data.

Next, periodogram method with the help from Hanning window was used to acquire the power spectral density (dB/rad/sample) plot of the raw data [18]. Using MATLAB, the average power spectral density plot was obtained with a suitable cut off frequency. The non-scintillation effects will be included if the cut off frequency is set to be too low while the scintillation effects in the received signal level are eliminated if the cut off frequency is set to be too high [16]. Hence, the cut off frequency needs to be cautiously set. Figure 4 depicts the average power spectral density. Based on Figure 4, a flat portion between 0.03 Hz to 0.1 Hz is seen and followed by a slope of approximately $f^{-8/3}$ above 0.1 Hz. For fully generated turbulence, the scintillation spectrum has a common pattern of low pass filtered white noise with a $-8/3$ slope [9].

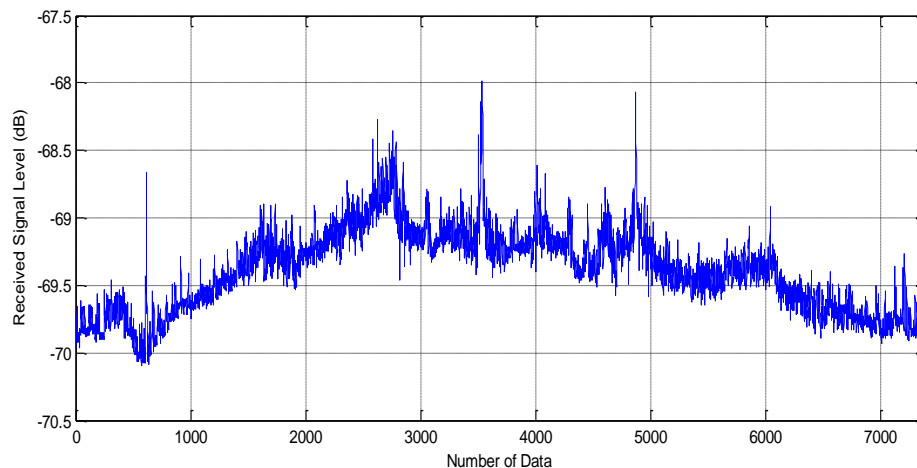


Figure 3. Measured raw data for 22 February 2016

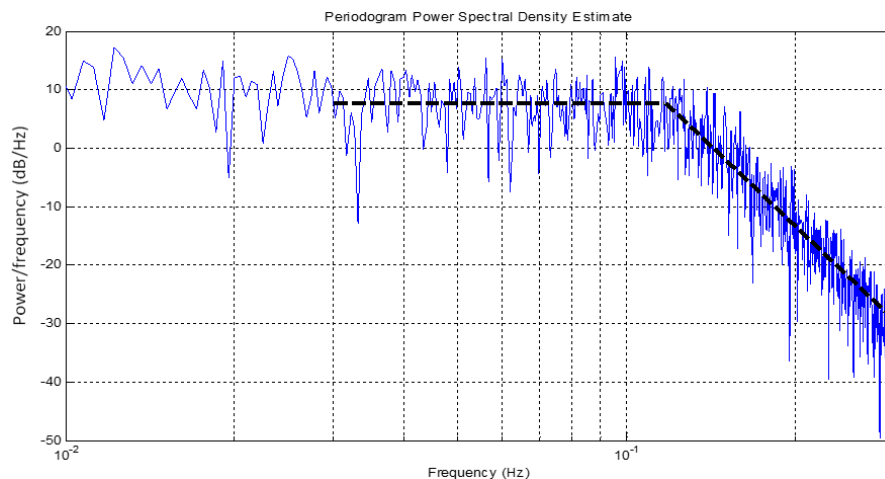


Figure 4. Average power spectral density

4.2. Analysis of scintillation data using MATLAB

Scintillation fades and enhancements were computed by using the cut off frequency of 0.11 Hz with the aid of high pass filter of order 10 so that the scintillation signals could be observed clearly [7]. The cut off frequency was obtained by calculating the average cut off frequency for the 12 months of year 2016 scintillation data as shown in Table 3. In MATLAB, after going through the filtering process, scintillation amplitudes for each day were displayed. After the scintillation amplitudes were obtained in the MATLAB, the amplitudes were then divided into two: scintillation fades and scintillation enhancements. Scintillation enhancements refer to the positive signal level of the scintillation amplitude whereas scintillation fades refer to the negative signal level of the scintillation amplitude. Scintillation enhancements produce an increase in intermodulation disturbance in a satellite transponder utilized in multicarrier and scintillation fades usually influence the earth station tracking systems, considering the usage of time constant in the system [18]. Figure 5 illustrates the scintillation amplitude for 13 April 2016.

Table 3. Cut off frequencies

Month	Cut-off Frequency (Hz)
January 2016	0.11
February 2016	0.12
March 2016	0.11
April 2016	0.12
May 2016	0.11
June 2016	0.12
July 2016	0.12
August 2016	0.11
September 2016	0.11
October 2016	0.11
November 2016	0.11
December 2016	0.12

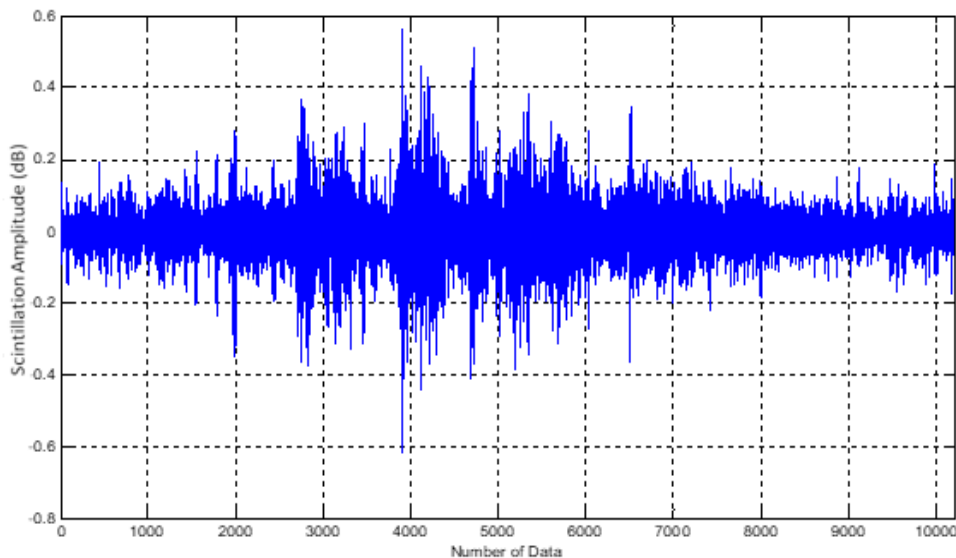


Figure 5. Scintillation amplitude of 13 April 2016

4.3. Data processing for cumulative distribution function (CDF)

Scintillation data were categorized into two parts: positive signal levels, scintillation enhancements and the negative signal levels, scintillation fades [19-25]. Then, the cumulative distribution was computed for scintillation fades. Figure 6 depicts the measured cumulative distribution function for scintillation fades on monthly basis. From Figure 6 it can be concluded that the measured scintillation fades in February 2016 are the highest until at 0.3% of the data, scintillation fades in March exceeded the rest of the months. Moreover, severe fading can be seen in February 2016 at 0.01% and followed by September 2016. Meanwhile, in November 2016, scintillation fades are the weakest.

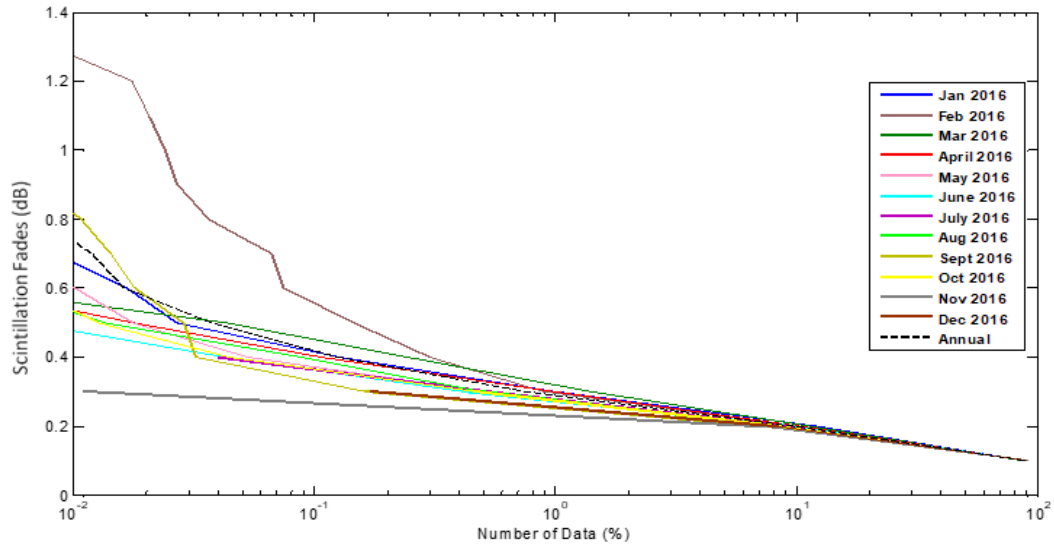


Figure 6. Measured monthly cumulative distribution for scintillation fades

The scintillation fades for all the months in 2016 are merely the same, but those plots branched out at approximately 10% of the data. In September 2016, the scintillation jumped further up till it reached its maximum at 0.8 dB and overlapped with scintillation fades in December 2016 at 0.17%. Besides that, the measured average annual cumulative distribution of scintillation fades exceeded the individual month at approximately at 1% of the data. However, in February 2016, the plot over predicted the average annual plot at every percentage number of data whereas in March 2016, the plot exceeded the average annual plot at 0.03% of the data. The values of the scintillation fades for each month and the percentage number of data are tabulated in Table 4.

Table 4. Measured monthly distributions of scintillation fades

Scintillation Fades (dB)	Number of Data (%)											
	Jan	Feb	March	April	May	June	July	Aug	Sept	Oct	Nov	Dec
0.1	86.5	87.88	86.7	88.5	90.2	89.8	90.2	91.1	92.07	90.7	92.6	91.4
0.2	12.33	10.52	11.64	10.4	9.18	9.78	9.25	8.3	7.67	8.74	7.36	8.43
0.3	0.97	0.87	1.37	0.96	0.48	0.40	0.51	0.48	0.17	0.45	0.01	0.17
0.4	0.13	0.3	0.24	0.11	0.05	0.04	0.04	0.087	0.03	0.04		
0.5	0.03	0.15	0.04	0.018	0.018	0.006		0.013	0.028	0.01		
0.6	0.02	0.075	0.0036	0.0032	0.01	0.003		0.0045	0.018	0.0063		
0.7	0.008	0.067			0.008				0.014			
0.8	0.006	0.036			0.005				0.01			
0.9	0.0021	0.027			0.003				0.007			
1.0		0.024							0.004			

4.4. Comparison of measured data with the existing scintillation models

The four models: Karasawa, ITU-R, Van De Kamp, and OTUNG are compared against the measured cumulative distribution. The comparison according to the location, elevation angle, frequency and data sampling are tabulated as in Table 5. All the cumulative distribution of the prediction models were computed using the equations elaborated in the literature review. From Figure 7 it can be deduced that the annual measured plot exceeds the scintillation fades for Karasawa. Similarly, from Figure 7, the scintillation fades for ITU-R is under predicted the annual measured plot at all percentages of time. Next, to compute the scintillation fades for Van de Kamp, cloud information was included. Based on Figure 7, the annual measured plot exceeds the scintillation fades for ITU-R model whereas for OTUNG, the scintillation fades over predicted the annual measured plot. However, at 0.3%, the annual measured plot exceeds the OTUNG scintillation fades plot. For this research study, the measured curve is compared with the other scintillation models to see whether the scintillation amplitude is comparable or not. The future work for this study is that the measured curve is modified according to the chosen scintillation models.

Table 5. Prediction models comparison

Research	Location	Elevation Angle (°)	Type of Antenna / Diameter (m)	Frequency (GHz)	Data Sampling (s)	Satellite
Karasawa [21]	Yamaguchi, Japan	4-30	Dish Antenna / 7.6	11.452	1	Intelsat-V
ITU-R [14]	Global	>4	Dish Antenna / 3-36	7-14	NA	NA
Van de Kamp [22]	Global	3-33	Dish Antenna	7-30	0.05	Satellite Links in Europe, Japan and US
OTUNG [17]	Sparsholt, UK	28.74	Diamond Shaped Antenna / 1.2	19.7704	0.1	Olympus
USM [7]	Penang, Malaysia	40.1	Cassegrain Antenna / 2.4	12.255	1.0	Superbird C
IIUM [11-12]	Gombak, Kuala Lumpur, Malaysia	77.5	Dish Antenna / 2.4	10.982	0.1	MEASAT-3
ASTRO (Measured)	Bukit Jalil, Kuala Lumpur, Malaysia	77.5	Dish Antenna / 13	12.202	0.05	MEASAT-3B

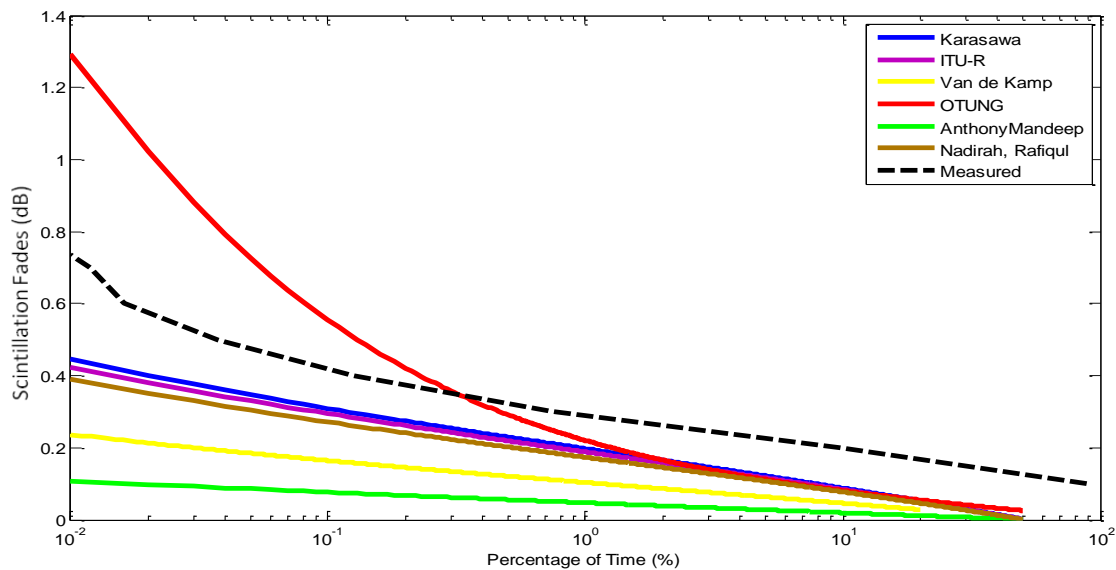


Figure 7. Scintillation fades of prediction model and measured data

5. CONCLUSION

In conclusion, the satellite communication study pertaining to the scintillation modeling in Ku-band has been reported and compared with the previous scintillation models. The scintillation effects on high elevation angle and high frequency have been analyzed and discussed. The results of the measured scintillation data from January 2016 till December 2016 are presented in this paper. Data from ASTRO with frequency of 12.202 GHz and elevation angle of 77.5° were used to observe the scintillation effects in the tropical climate. The raw data were processed in the MATLAB to gain a suitable cut off frequency. Then, the average cut off frequency was used in the high pass filter of order 10 to eliminate any propagation factors that lead to other propagation effects. The scintillation data were then filtered and divided into two: scintillation fades and scintillation enhancements. However, this paper is only focused on the scintillation fades and therefore the data for scintillation enhancements were ignored. The cumulative distributions of scintillation fades were computed based on monthly and annual. Then, the annual cumulative data were compared with the six existing scintillation fades model namely Karasawa, ITU-R, OTUNG, Van de Kamp, Anthony-Mandeeep, and Nadirah-Rafiqul.

Based on the finding, the annual measured scintillation fades over predicted all other existing scintillation models which has the highest scintillation fades reading. The scintillation fades at 0.01% for annual measured data is 0.73 dB whereas for OTUNG is 0.55 dB, Karasawa is 0.45 dB, ITU-R is 0.42 dB, Nadirah-Rafiqul is 0.4 dB, Van de Kamp is 0.42 dB, and Anthony-Mandeeep 0.11 dB. Therefore, the objectives of this research study are met accordingly, which are: to analyze the scintillation data obtained for Ku-band and to compare and validate the results with other scintillation models.

This research can be improved by analyzing the scintillation enhancements for the measured data and compare it with the existing scintillation enhancements models. It is also recommended to study the monthly, seasonal, diurnal and annual statistics of both scintillation enhancements and scintillation fades. In order to know how reliable, the model is for this measured data, a new scintillation model based on the measured data need to be computed. Therefore, more research activities should be carried out after comparing the annual measured cumulative distribution with the other existing models.

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